

Preprint typeset in JHEP style. - PAPER VERSION

LUTP/01/03, DCTP/01/06, IPPP/01/03, hep-ph/0101226

Resumming the Light Hemisphere Mass and Narrow Jet Broadening distributions in e^+e^- annihilation *

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ABSTRACT: We present predictions of two event shape distributions, the light hemisphere mass and the narrow jet broadening, to next-to-next-to-leading logarithmic order. We apply the coherent branching formalism to resum the leading $\mathcal{O}(\alpha_s^n L^{2n-1})$ and sub-leading infrared logarithms down to $\mathcal{O}(\alpha_s^n L^{2n-3})$ to all orders in the coupling constant. We include the recently calculated non-logarithmic next-to-leading order contributions. Applying simple power corrections to the resummed result gives good agreement with the available data from LEP.

KEYWORDS: QCD, Jets, LEP HERA and SLC Physics, NLO and NNLO Computations.

*Work supported in part by the UK Particle Physics and Astronomy Research Council and by the EU Fourth Framework Programme ‘Training and Mobility of Researchers’, Network ‘Quantum Chromodynamics and the Deep Structure of Elementary Particles’, contract FMRX-CT98-0194 (DG 12 - MHT).

1. Introduction

Over the last decade, event shape distributions in e^+e^- annihilation have provided some of the most precise tests of QCD. Most of the attention has focussed on three-jet event shape variables such as thrust, wide jet broadening, heavy hemisphere mass, C parameter etc.. These variables may be generated by single gluon emission from the underlying $q\bar{q}$ pair in the e^+e^- annihilation process which first contributes at $\mathcal{O}(\alpha_s)$. The analysis of such events is rather sophisticated. In the first instance, fixed order QCD perturbation theory [1] is employed (at next-to-leading order (NLO)). However, when the observable O is small and $L = \ln(1/O)$ is large the distribution suffers large logarithmic contributions of the form $\mathcal{O}(\alpha_s^n L^{2n-1})$ that render the fixed order perturbation theory insufficient. In many cases the dominant infrared logarithms can be resummed using the coherent branching formalism [2, 3, 4, 5]. Furthermore, significant power suppressed effects are present and have been phenomenologically studied [6]. To accurately describe the data, all of these contributions – fixed order, infrared resummation and power correction – are found to be necessary. Perhaps the most important measurement extracted from three-jet event shape data is the determination of the strong coupling constant.

Four-jet event shape observables also contain useful information about QCD. They are more sensitive to the triple gluon vertex and therefore the true gauge structure of QCD [7] and also to the presence of other light coloured particles such as the gluino [8] that can be pair produced by gluon splitting. However, these variables have received much less attention partly because they are suppressed by an additional power of α_s requiring a second gluon to be radiated but also because the theoretical description is much less developed. Recently however, four separate general purpose Monte Carlo programs have been developed to estimate the next-to-leading order, $\mathcal{O}(\alpha_s^3)$, corrections to four-jet event shapes: **MENLO PARC** [9], **DEBRECEN** [10] and **MERCUTIO** [11] employing on the one-loop helicity amplitudes for $e^+e^- \rightarrow 4$ partons [12] and **EERAD2** [13] based on the interference of the one-loop matrix element with tree level [14]. For most four-jet event shape variables, the situation is the same as for the three-jet event shape variables; the NLO corrections are very large and, for renormalisation scales of the order of the centre-of-mass energy Q , still undershoot the data significantly [13]. This points at the presence of large infrared effects as well as significant power corrections. With the exception of the four-jet rate in certain schemes [15] and near-to-planar three-jet events [16], the issue of resumming infrared logarithms for specific four-jet event shape variables has not yet been addressed. Similarly, power suppressed effects have only been studied for the out-of-event plane momentum distribution [16]. It is the goal of this Letter to resum the leading $\mathcal{O}(\alpha_s^n L^{2n-1})$ and sub-leading infrared logarithms down to $\mathcal{O}(\alpha_s^n L^{2n-3})$ to all orders in the coupling constant for the light hemisphere mass ρ_L and narrow jet broadening B_N distributions.

To be more precise let us first define these variables. We first separate the event at centre-of-mass energy $Q = \sqrt{s}$ into two hemispheres H_1, H_2 divided by the plane normal to the thrust axis \mathbf{n}_T . Particles that satisfy $\mathbf{p}_i \cdot \mathbf{n}_T > 0$ are assigned to hemisphere H_1 , while all other particles are in H_2 . Jet broadening measures the summed scalar momentum transverse to the thrust axis in one of the hemispheres while the hemisphere mass is the invariant mass of the hemisphere,

$$B_N = \min_{i=1,2} \frac{\sum_{\mathbf{p}_k \in H_i} |\mathbf{p}_k \times \mathbf{n}_T|}{2 \sum_k |\mathbf{p}_k|} \quad (1.1)$$

$$\rho_L = \frac{1}{s} \cdot \min_{i=1,2} \left(\sum_{\mathbf{p}_k \in H_i} p_k \right)^2. \quad (1.2)$$

Note that this definition of the light hemisphere mass is the common modification of the original variables suggested by Clavelli [18]. These four-jet event shape variables are intimately connected to their three-jet event shape counterparts, the wide jet broadening and heavy hemisphere mass,

$$B_W = \max_{i=1,2} \frac{\sum_{\mathbf{p}_k \in H_i} |\mathbf{p}_k \times \mathbf{n}_T|}{2 \sum_k |\mathbf{p}_k|} \quad (1.3)$$

$$\rho_H = \frac{1}{s} \cdot \max_{i=1,2} \left(\sum_{\mathbf{p}_k \in H_i} p_k \right)^2 \quad (1.4)$$

that have the property of exponentiation. That is to say that the fraction of events where the observable $O = B_W$ or $O = \rho_H$ has a value less than O obeys the exponentiated form, [4, 5]

$$R(O, \alpha_s(Q^2)) = \int_0^O \frac{1}{\sigma} \frac{d\sigma}{dO} dO \quad (1.5)$$

$$= C(\alpha_s(Q^2)) \Sigma(O, \alpha_s(Q^2)) + D(O, \alpha_s(Q^2)) \quad (1.6)$$

where

$$C(\alpha_s(Q^2)) = 1 + \sum_{n=1}^{\infty} C_n \alpha_s^n \quad (1.7)$$

$$\begin{aligned} \ln \left(\Sigma(O, \alpha_s(Q^2)) \right) &= \sum_{n=1}^{\infty} \sum_{m=1}^{n+1} G_{nm} \alpha_s^n L^m \\ &= L g_1(\alpha_s L) + g_2(\alpha_s L) + \alpha_s g_3(\alpha_s L) + \dots \end{aligned} \quad (1.8)$$

$$D(O, \alpha_s(Q^2)) = \sum_{n=0}^{\infty} D_n \alpha_s^n. \quad (1.9)$$

Here C_n and G_{nm} are constants and the perturbatively calculable coefficients $D_n \rightarrow 0$ as $O \rightarrow 0$. Knowledge of g_2 (or equivalently all G_{nn}) allows resummation of terms in

Σ down to $\mathcal{O}(\alpha_s^n L^n)$. The accuracy of R depends on the number of terms known in $C(\alpha_s(Q^2))$, with each known term giving information about two towers of logarithms. For example, $C = 1$ resums the $\mathcal{O}(\alpha_s^n L^{2n})$ and $\mathcal{O}(\alpha_s^n L^{2n-1})$ terms in R , C_1 resums the $\mathcal{O}(\alpha_s^n L^{2n-2})$ and $\mathcal{O}(\alpha_s^n L^{2n-3})$ terms and so on.

2. Coherent branching

The coherent branching formalism allows the resummation of soft and collinear logarithms due to the emission of gluons from a hard parton. As a specific example, let us consider the jet mass distribution $J^a(Q, k^2)$ as the probability of producing a final state jet with invariant mass k^2 from a parent parton a produced in a hard process at the scale Q^2 . For an initial quark this is [3]

$$J^q(Q^2, k^2) = \delta(k^2) + \int_0^{Q^2} \frac{d\tilde{q}^2}{\tilde{q}^2} \int_0^1 dz \Theta(z^2(1-z)^2\tilde{q}^2 - Q_0^2) P^{qq}[\alpha_s(z^2(1-z)^2\tilde{q}^2), z] \\ \times \left[\int_0^\infty dq^2 \int_0^\infty dk'^2 \delta\left(k^2 - z(1-z)\tilde{q}^2 - \frac{k'^2}{z} - \frac{q^2}{1-z}\right) J^q(z^2\tilde{q}^2, k'^2) J^g((1-z)^2\tilde{q}^2, q^2) \right. \\ \left. - J^q(\tilde{q}^2, k^2) \right] \quad (2.1)$$

normalised such that

$$\int_0^\infty dk^2 J^a(Q^2, k^2) = 1 \quad (2.2)$$

and where the next-to-leading order $q \rightarrow qg$ splitting kernel in the \overline{MS} scheme with N_f flavours is

$$P_{qq}[\alpha_s, z] = \frac{\alpha_s}{2\pi} C_F \frac{1+z^2}{1-z} \left(1 + \frac{\alpha_s}{2\pi} K\right) + \dots, \quad (2.3)$$

where

$$K = C_A \left(\frac{67}{18} - \frac{\pi^2}{6} \right) - \frac{5}{9} N_f. \quad (2.4)$$

Eq. (2.1) has a simple physical interpretation. The first term is the possibility that the originating parton does not emit any radiation. The transverse momentum of the parton is therefore unchanged. Alternatively, the quark may branch into a quark and a gluon subject to the phase space constraints of two body decay and which subsequently undergo further emissions. This is described by the term proportional to $J^q J^g$. The last term is due to virtual corrections and ensures that soft and

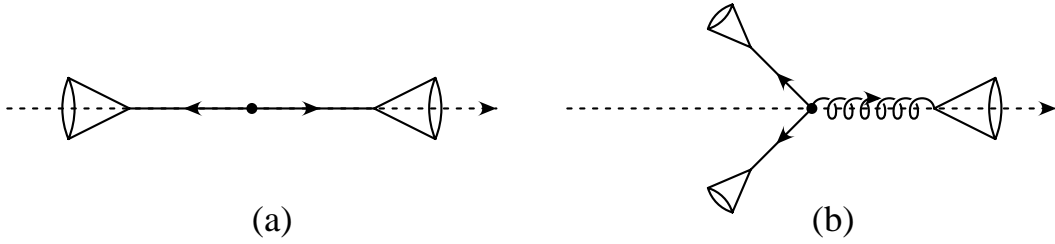


Figure 1: Event pictures of (a) two-jet configurations from quark-antiquark final states and (b) three-jet configurations originating from $q\bar{q}g$ events where the gluon is the hardest parton. The cones represent coherent soft and collinear gluon emission. The thrust axis is denoted by a dashed line.

collinear singularities are regularised. A similar equation holds for J^g and involves the $g \rightarrow gg$ and $g \rightarrow q\bar{q}$ splitting kernels. Solving the integral equation for J^a is equivalent to resumming the infrared logarithms. In particular, the probability that an isolated parton a forms a jet with mass less than ρQ^2 is given by

$$\Sigma^a(\rho, \alpha_s(Q^2)) = \int_0^{\rho Q^2} dk^2 J^a(Q^2, k^2). \quad (2.5)$$

In Ref. [3], $\ln(\Sigma_H^a)$ is solved to next-to-leading logarithmic accuracy and Σ_H^a is therefore known to $\mathcal{O}(\alpha_s^n L^n)$.

Similarly we can define the function $T^a(Q, \mathbf{k}_t; p_t)$ [5] which describes the distribution of the summed scalar transverse momentum p_t in a jet of parton type a produced with vector transverse momentum \mathbf{k}_t , at scale Q . The structure is identical to Eq. (2.1).

3. The probabilistic interpretation

We can apply the coherent branching formalism to event shapes in the following schematic way illustrated in Fig. 1. The underlying configuration in e^+e^- annihilation is the production of a quark-antiquark pair aligned with the thrust axis (Fig. 1(a)). A fraction r_2 of the events have this structure where

$$r_2 = 1 + \mathcal{O}(\alpha_s). \quad (3.1)$$

Each parton then undergoes soft and collinear gluon emission (denoted by the open cone at the head of the parton). This contribution describes small angle and soft emission accurately in *two-jet-like* events when B_W and ρ_H are small and gives rise to the exponentiated first term in Eq. (1.6). However, it does not describe the possibility of wide angle gluon emission shown in Fig. 1(b) where the gluon is the

hardest parton. This *three-jet-like* matrix element correction is not logarithmically enhanced at $\mathcal{O}(\alpha_s)$ and produces a fraction r_3 of events given by

$$R_L^{(3)} = C_F \frac{\alpha_s}{2\pi} \int dx_g \left[\frac{4 - 4x_g + 4x_g^2}{x_g} \ln \left(\frac{2x_g - 1}{1 - x_g} \right) - 6x_g + 4 \right] \mathcal{S}(x_g Q) + \mathcal{O}(\alpha_s^2). \quad (3.2)$$

The function $\mathcal{S}(x_g Q)$ represents the subsequent showering of the hard gluon. This will have a non-trivial dependence on the energy fraction x_g , setting the initial scale for the branching. In this configuration the gluon energy fraction will be in the region $\frac{2}{3} \leq x_g \leq 1$. Taking the approximation $x_g = 1$ inside \mathcal{S} enables the factorisation of the fraction of events, r_3 ,

$$R_L^{(3)} = r_3 \mathcal{S}(Q) + \mathcal{O}(\alpha_s^2), \quad (3.3)$$

where at $\mathcal{O}(\alpha_s)$

$$\begin{aligned} r_3 &= \frac{\alpha_s}{2\pi} C_F \left(2 \ln(2)^2 - \frac{5}{4} \ln(3) + 4 \text{Li}_2 \left(-\frac{1}{2} \right) + \frac{\pi^2}{3} - \frac{1}{6} \right) \\ &\sim \frac{\alpha_s}{2\pi} C_F (0.917). \end{aligned} \quad (3.4)$$

The thrust axis is now aligned with the gluon and the values of B_W and ρ_H are determined by the energies and angular separation of the quark and antiquark. In general B_W and ρ_H will not be small and in fact, the phase space for small B_W and ρ_H configurations is vanishingly small. Therefore, for these observables we can associate the three-jet-like contribution with $D(O, \alpha_s(Q^2))$ in Eq. (1.6). Only the two-jet configuration is logarithmically enhanced and as discussed earlier, the coherent branching algorithm has been used [3, 4, 5] to resum logarithmic terms in the perturbative expansion of Σ down to $\mathcal{O}(\alpha_s^n L^n)^*$.

On the other hand, the three jet-like configuration is *not* suppressed for small values of the four-jet variables B_N and ρ_L which are generated by subsequent branching of the gluon. This type of branching engenders leading perturbative contributions to R_L and R_N of $\mathcal{O}(C_F C_A^{n-1} \alpha_s^n L^{2n-2})$ which can be resummed. However in approximating the hard scale at the initiation of the gluon shower, we have spoiled the next-to-leading logarithm terms. This effect will introduce an uncertainty in the fourth tower of logarithms, i.e. at $\mathcal{O}(\alpha_s^n L^{2n-3})$. In principle, this loss of accuracy can be avoided by retaining the gluon energy dependence in Eq. (3.2). However, as

*We note that in solving the integral equations for the jet functions, certain approximations may need to be made that may limit the number of logarithmic terms that are actually resummed. For example, if the coefficient G_{22} is not determined accurately, then terms of $\mathcal{O}(\alpha_s^n L^{2n-2})$ are not exactly resummed. Dokshitzer et al [19] have recently re-studied the wide jet broadening and found that in order to obtain G_{22} correctly requires a careful treatment of quark recoil.

we will see later, this contribution turns out to be numerically negligible, thereby justifying the approximation.

Let us first focus on the hemisphere masses. The fraction of events with heavy hemisphere mass less than ρ_H i.e. $k_1^2 < \rho_H Q^2$ and $k_2^2 < \rho_H Q^2$ is given entirely by the two jet contribution (with $r_2 = 1 + \mathcal{O}(\alpha_s)$)

$$R_H(\rho_H, \alpha_s(Q^2)) \underset{\rho_H \ll 1}{=} r_2 \left(\int_0^\infty dk_1^2 \int_0^\infty dk_2^2 J^q(Q^2, k_1^2) J^q(Q^2, k_2^2) \Theta(\rho_H Q^2 - k_1^2) \Theta(k_1^2 - k_2^2) \right. \\ \left. + \int_0^\infty dk_1^2 \int_0^\infty dk_2^2 J^q(Q^2, k_1^2) J^q(Q^2, k_2^2) \Theta(\rho_H Q^2 - k_2^2) \Theta(k_2^2 - k_1^2) \right), \quad (3.5)$$

and where the constraint that $\rho_H \ll 1$ has suppressed the three and more jet contributions. Following the steps of Ref. [5], we can rewrite this formula using the phase space restrictions as

$$R_H(\rho_H, \alpha_s(Q^2)) \underset{\rho_H \ll 1}{=} r_2 \left[\Sigma_H^q(\rho_H, \alpha_s(Q^2)) \right]^2. \quad (3.6)$$

where

$$\Sigma_H^q(\rho_H, \alpha_s(Q^2)) = \int_0^{\rho_H Q^2} dk^2 J^q(Q^2, k^2). \quad (3.7)$$

The $\mathcal{O}(\alpha_s)$ contribution to r_2 is fixed by requiring that when ρ_H reaches its maximum value of $1/3$ for three parton configurations, $R_H = 1 + \mathcal{O}(\alpha_s)^2$.

On the other hand, the fraction of events with light hemisphere mass less than ρ_L receives contributions from both two and three-jet configurations. For the two jet case, $k_1^2 < \rho_L Q^2$ and $k_2^2 > k_1^2$ and vice versa while the three-jet contribution arises when $k_g^2 < \rho_L Q^2$. In the small ρ_L limit, what happens in the quark-antiquark hemisphere is irrelevant and k_1^2 and k_2^2 are unbounded. Altogether we have,

$$R_L(\rho_L, \alpha_s(Q^2)) \underset{\rho_L \ll 1}{=} r_2 \left(\int_0^\infty dk_1^2 \int_0^\infty dk_2^2 J^q(Q^2, k_1^2) J^q(Q^2, k_2^2) \Theta(\rho_L Q^2 - k_1^2) \Theta(k_2^2 - k_1^2) \right. \\ \left. + \int_0^\infty dk_1^2 \int_0^\infty dk_2^2 J^q(Q^2, k_1^2) J^q(Q^2, k_2^2) \Theta(\rho_L Q^2 - k_2^2) \Theta(k_1^2 - k_2^2) \right) \\ + r_3 \int_0^\infty dk_g^2 J^g(Q^2, k_g^2) \Theta(\rho_L Q^2 - k_g^2) \int_0^\infty dk_1^2 J^q(Q^2, k_1^2) \int_0^\infty dk_2^2 J^q(Q^2, k_2^2) \quad (3.8)$$

where to $\mathcal{O}(\alpha_s^2)$ the two jet fraction r_2 is given by,

$$r_2 = 1 - r_3 + \mathcal{O}(\alpha_s^2). \quad (3.9)$$

Simplifying the phase space constraints and utilising the normalisation condition (2.5), we find

$$R_L(\rho_L, \alpha_s(Q^2)) \underset{\rho_L \ll 1}{=} r_2 \left(2 \Sigma_H^q(\rho_L, \alpha_s(Q^2)) - [\Sigma_H^q(\rho_L, \alpha_s(Q^2))]^2 \right) + r_3 \Sigma_H^g(\rho_L, \alpha_s(Q^2)), \quad (3.10)$$

where Σ_H^g is defined as an integral over the gluon jet mass distribution J^g in a similar way to Eq. (3.7). The functions that resum the logarithms for the light hemisphere mass are the same as those that resum the logarithms for the heavy hemisphere mass. Now however, exponentiation in its purest form is spoiled because the final result is a sum of terms.

The analysis for the narrow jet broadening proceeds in the same way. We find that R_N , the probability of finding an event with a light hemisphere mass less than B_N , is given by

$$R_N(B_N, \alpha_s(Q^2)) \underset{B_N \ll 1}{=} r_2 \left(2 \Sigma_W^q(B_N, \alpha_s(Q^2)) - [\Sigma_W^q(B_N, \alpha_s(Q^2))]^2 \right) + r_3 \Sigma_W^g(B_N, \alpha_s(Q^2)), \quad (3.11)$$

where the probability of obtaining a jet with summed scalar transverse momentum p_t with respect to the jet axis less than $2BQ$ starting from a parton of type a , Σ_W^a , is given by,

$$\Sigma_W^a(B, \alpha_s(Q^2)) = \int_0^{2BQ} T^a(Q, \mathbf{0}, p_t) dp_t. \quad (3.12)$$

4. All-orders resummation of large logarithms

In this section we discuss the all-orders resummation of leading logarithms $\mathcal{O}(\alpha_s^n L^{2n})$ as well as sub-leading logarithms down to $\mathcal{O}(\alpha_s^n L^{2n-2})$. As discussed in the previous section, a resummation of this order is achieved by considering both two-jet and three-jet configurations. From eqs. (3.10) and (3.11), we see that to determine R_L and R_N requires knowledge of Σ_H^a and Σ_W^a respectively. Both of these functions have the exponentiated form (1.6) and the exponents have been solved to single logarithmic accuracy. Σ_H^a is therefore known to $\mathcal{O}(\alpha_s^n L^n)$.

Explicit expressions for Σ_H valid to this order are given in [3] and, introducing the renormalisation scale dependence in the standard manner and dropping the parton

index, we reproduce them here for illustrative purposes,

$$\begin{aligned}\Sigma_H(\rho, \alpha_s(\mu^2), \frac{Q^2}{\mu^2}) &= \frac{\exp[\mathcal{F}(\alpha_s(\mu^2), L)]}{[1 - \mathcal{S}(\alpha_s(\mu^2), L)]} \\ &= \frac{\exp[Lf_1(x) + f_2(x) + x^2 f_1'(x) \ln(\mu^2/Q^2)]}{[1 - f_1(x) - x f_1'(x)]} + \mathcal{O}(\alpha_s^n L^{n-1})\end{aligned}\quad (4.1)$$

where

$$L = \ln(1/\rho), \quad x = \beta_0 \alpha_s(\mu^2) L. \quad (4.2)$$

The functions f_1 , f_2 and f_1' are

$$f_1(x) = -\frac{A^{(1)}}{2\pi\beta_0 x} \left[(1-2x) \ln(1-2x) - 2(1-x) \ln(1-x) \right], \quad (4.3)$$

$$\begin{aligned}f_2(x) &= -\frac{A^{(2)}}{2\pi^2\beta_0^2} \left[2 \ln(1-x) - \ln(1-2x) \right] \\ &\quad + \frac{B^{(1)}}{2\pi\beta_0} \ln(1-x) - \frac{A^{(1)}\gamma_E}{\pi\beta_0} \left[\ln(1-x) - \ln(1-2x) \right] \\ &\quad - \frac{A^{(1)}\beta_1}{2\pi\beta_0^3} \left[\ln(1-2x) - 2 \ln(1-x) + \frac{1}{2} \ln^2(1-2x) - \ln^2(1-x) \right],\end{aligned}\quad (4.4)$$

$$f_1'(x) = \frac{A^{(1)}}{2\pi\beta_0 x^2} \left[\ln(1-2x) - 2 \ln(1-x) \right], \quad (4.5)$$

with

$$\beta_0 = \frac{11C_A - 2N_f}{12\pi}, \quad \beta_1 = \frac{17C_A^2 - 5C_A N_f - 3C_F N_f}{24\pi^2}, \quad (4.6)$$

For quarks,

$$A^{(1)} = C_F, \quad A^{(2)} = \frac{1}{2} C_F K, \quad B^{(1)} = -\frac{3}{2} C_F, \quad (4.7)$$

while for gluons,

$$A^{(1)} = C_A, \quad A^{(2)} = \frac{1}{2} C_A K, \quad B^{(1)} = -2\pi\beta_0, \quad (4.8)$$

with K given by Eq.(2.4).

Altogether Eqs. (4.1) to (4.8) are sufficient to determine Σ_H^q and Σ_H^g to $\mathcal{O}(\alpha_s^n L^n)$. However, this does not determine R_N of Eq. 3.11 to the same order due to the approximations made in categorising two and three-jet configurations before showering. In separating these events, we have introduced an error in the resummation of the

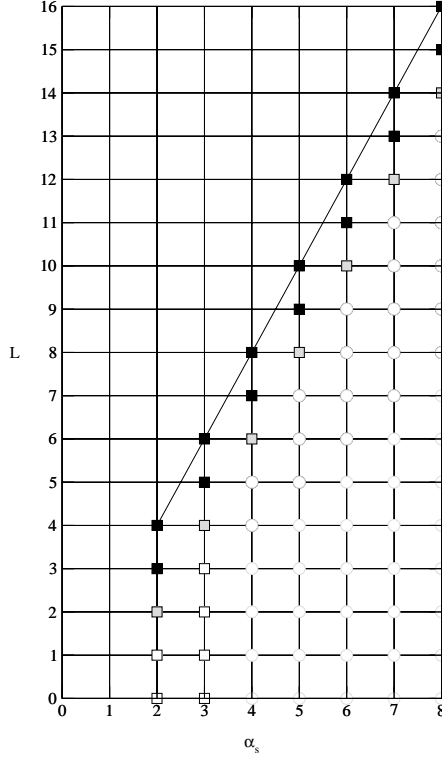


Figure 2: The towers of infrared logarithms appearing in the four-jet event shape rates R_L and R_N . The resummation includes all terms down to $\mathcal{O}(\alpha_s^n L^{2n-2})$ (three towers) denoted by filled squares and complete $\mathcal{O}(\alpha_s^2, \alpha_s^3)$ contributions from fixed order denoted by the empty squares. All other terms are incomplete and denoted by empty circles. The black filled squares denote terms generated purely in the two-jet limit. The grey filled squares denote contributions from soft gluons showering off a hard gluon.

showering off the hard gluon at next-to-leading logarithm level. The expression for R_N therefore correctly sums to all orders in the strong coupling, α_s , only the leading three towers of large logarithms from $\mathcal{O}(\alpha_s^n L^{2n})$ down to $\mathcal{O}(\alpha_s^n L^{2n-2})$.

Because knowledge of Σ_H is currently limited to $\mathcal{O}(\alpha_s^n L^n)$, we can only correctly resum logarithms down to $2n - 2 \geq n$ or equivalently $n \geq 2$. The resummed formula (3.10) does not include the $\mathcal{O}(\alpha_s^2 L)$ term (just as the analogous formulae for resumming three jet variables do not include the $\mathcal{O}(\alpha_s^2 L)$ term) present in the lowest order perturbative coefficient. Similarly, it does not produce the $\mathcal{O}(\alpha_s^3 L^3)$ to $\mathcal{O}(\alpha_s^3 L)$ terms that occur in the next-to-leading order perturbative coefficient. The perturbative calculation provides the α_s^2 and α_s^3 contributions exactly and therefore the most significant omitted term is $\mathcal{O}(\alpha_s^4 L^5)$, see Fig. 2.

Precisely the same discussion applies to the narrow jet broadening. Using the coherent branching formalism Σ_W has been determined to $\mathcal{O}(\alpha_s^n L^n)$ accuracy and

Catani et al (CTW) have provided analogous expressions for Σ_W that are given in [5]. However, in doing so certain simplifying approximations concerning the recoil transverse momentum have been made. As a consequence, the terms of $\mathcal{O}(\alpha_s^n L^n)$ are incomplete. Dokshitzer and collaborators [19] have found that treating the quark recoil more carefully causes the CTW result for Σ_W to be adjusted by a multiplicative factor. This generates the correct $\alpha_s^n L^n$ terms. We do not give the final form for Σ_W here, but instead refer the interested reader to Ref. [19]. Inserting the recoil-corrected form for Σ_W in the resummed expression for R_N (3.11) again allows resummation of the first three towers of logarithms, i.e. down to $\mathcal{O}(\alpha_s^n L^{2n-2})$, as shown in Fig. 2.

5. Numerical results

As usual, the resummed result contains part of the fixed order perturbative contribution and the overlap must be removed by matching. This is done by expanding the resummed result as a series in the strong coupling constant and explicitly removing the terms corresponding to the fixed order calculations. This can be achieved in several ways of which R matching and $\ln(R)$ matching are the most common. In the R matching scheme, the coefficients of each of the unsummed logarithms present in the fixed order perturbative coefficients must be numerically extracted. For the four-jet event shape observables discussed here, this corresponds to determining the coefficient of $\alpha_s^2 L$ from the lowest order perturbative contribution and the coefficients of $\alpha_s^3 L^2$ and $\alpha_s^3 L$ from the next-to-leading order contribution. This is impractical. However, in the more commonly used $\ln(R)$ matching scheme, it is assumed that the fixed order result exponentiates and therefore it is not necessary to make this extraction because any logarithmic terms remaining after subtracting the overlap from the fixed order contribution are exponentially suppressed. We therefore employ the $\ln(R)$ matching procedure.

We expect that at large values of the observable O the resummed result is dominated by the fixed order calculation. However, the resummed expressions (3.10) and (3.11) valid in the small ρ_L and B_N limits do not contain information about the kinematic endpoints of the distributions. To ensure that the resummed result vanishes at the endpoint we make the substitution

$$\frac{1}{O} \rightarrow \frac{1}{O} - \frac{1}{O^{\max}} + 1, \quad (5.1)$$

where O^{\max} corresponds to the endpoint of the distribution at the accuracy of the fixed order calculation. We use $\rho_L^{\max} = 0.167$ and $B_N^{\max} = 0.204$.

Numerical results for the light hemisphere mass and for the narrow jet broadening are shown in figures 3 and 4 respectively. Throughout we set $\mu = Q = M_Z$ and use $\alpha_s(M_Z) = 0.118$ corresponding to the current world average. The next-to-leading order result which diverges at small values of the event shapes is taken from [13] and

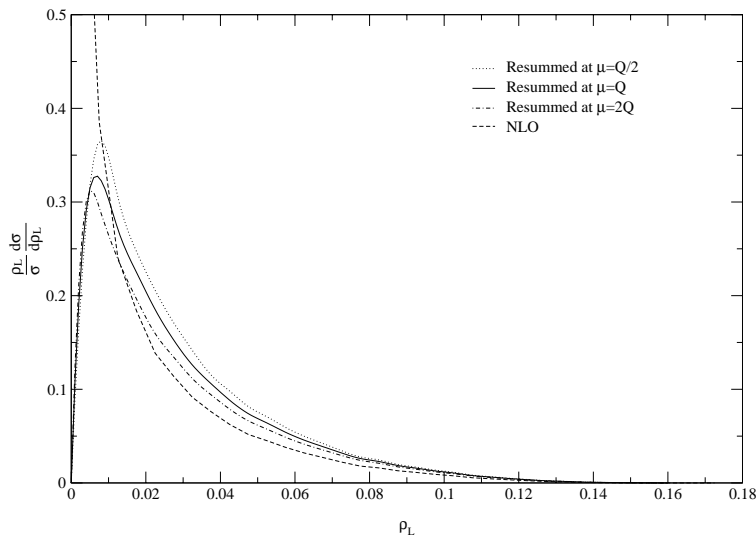


Figure 3: The resummed (solid) and fixed order NLO (dashed) predictions for the light hemisphere mass distribution $\frac{\rho_L}{\sigma} \frac{d\sigma}{d\rho_L}$ at $\mu = Q = M_Z$. The resummed prediction at $\mu = Q/2$ ($\mu = 2Q$) is shown as a dotted (dot-dashed) curve.

is evaluated at $\mu = Q$. Although formally, the three-jet like contribution is needed to resum terms in R_L and R_N of $\mathcal{O}(\alpha_s^n C_F C_A^{n-1} L^{2n-2})$ we find that it is numerically insignificant, mainly due to the smallness of r_3 . This justifies the earlier approximation of neglecting the gluon energy dependence in three-jet-like configuration.

Figures 3 and 4 show that the resummations are extremely important for $\rho_L < 0.01$ and $B_N < 0.02$. Rather than the divergent fixed order prediction, we have the more physical resummed result that the probability of finding events with no radiation (very small values of ρ_L and B_N) are vanishingly small. For the reference value of $\mu = Q$, the peak position of the ρ_L distribution occurs at $\rho_L = 0.01$ with a height of 0.33, while for B_N the peak occurs at $B_N = 0.02$ with a value of 0.48. At larger values, the resummation changes the NLO prediction by a more moderate amount indicating that uncalculated higher order corrections are under control. At very large values of O , the resummed and NLO predictions coincide because of the matching procedure of Eq. (5.1).

To illustrate the residual renormalisation scale dependence, we also show the effect of varying μ by a factor of 2 either side of the reference value $\mu = Q$. We see that around the peak region, different scale choices alter the prediction by $\pm 10\%$.

The infrared resummation significantly improves the perturbative prediction for the event shape observable. However, when comparing with experimental data we should be aware that important non-perturbative hadronisation corrections are present. The effect of hadronisation on the distribution is to shift the value of the

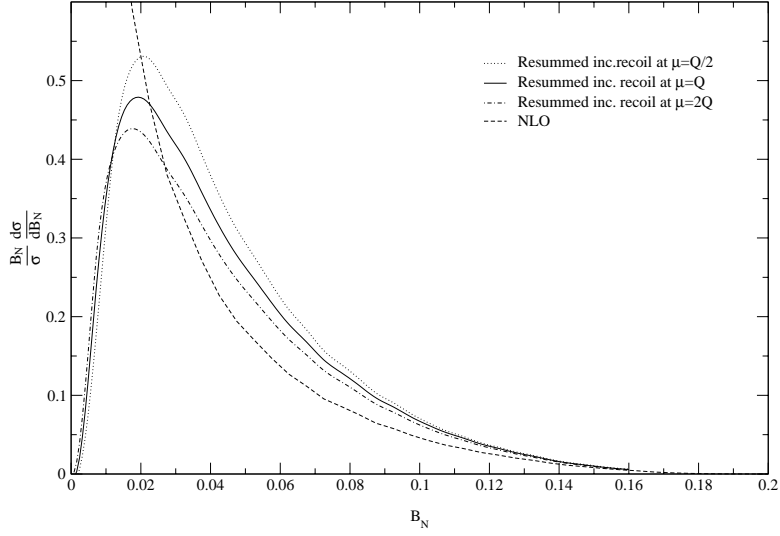


Figure 4: The resummed (solid) and fixed order NLO (dashed) predictions for the narrow jet broadening distribution $\frac{B_N}{\sigma} \frac{d\sigma}{dB_N}$ at $\mu = Q = M_Z$. The resummed prediction at $\mu = Q/2$ ($\mu = 2Q$) is shown as a dotted (dot-dashed) curve.

observable away from the two-jet region,

$$O \rightarrow O + O_{\text{NP}}, \quad (5.2)$$

where the non-perturbative correction depends on the typical hadron scale $\mathcal{O}(1 \text{ GeV})$ and is suppressed by a power of Q . In principle these power corrections can be estimated using the dispersive approach of Ref. [20] where a non-perturbative parameter μ_I is introduced to describe the running of α_s in the infrared region. For the associated three jet variables the non-perturbative corrections are typically estimated to be $\mathcal{O}(1 \text{ GeV}/Q)$ for ρ_H [20] and $\mathcal{O}(0.3 \text{ GeV} \ln(1/B_W)/Q)$ for B_W [21] and arise through the hadronisation of one of the two jets in the event. Because the four-jet event shapes are largely related to what happens in the second jet, we might expect that the hadronisation corrections are similar. To illustrate the potential effects of hadronisation in the four-jet event shapes, we just transfer these corrections directly so that,

$$\rho_L \rightarrow \rho_L + \frac{1 \text{ GeV}}{Q}, \quad (5.3)$$

$$B_N \rightarrow B_N + \frac{0.3 \text{ GeV} \ln(1/B_N)}{Q}. \quad (5.4)$$

The simplified hadronisation correction applied to the resummed distributions for the light hemisphere mass and narrow jet broadening with $\mu = Q$ is shown in figures 5 and 6 respectively. To emphasize the dramatic effect the power correction has, we also show the uncorrected predictions. There are two effects. First the distribution

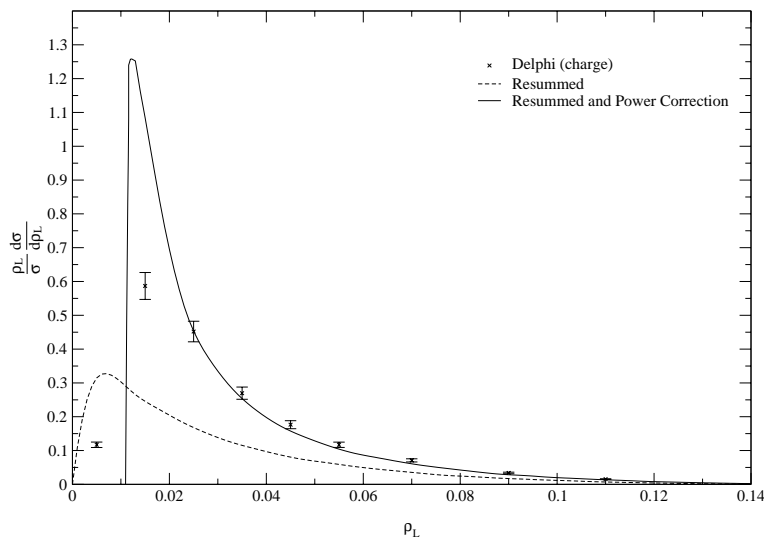


Figure 5: The resummed prediction for the light hemisphere mass distribution $\frac{\rho_L}{\sigma} \frac{d\sigma}{d\rho_L}$ at $\mu = Q = M_Z$ modified by a non-perturbative power correction $\rho_L \rightarrow \rho_L + 1 \text{ GeV}/Q$. The resummed prediction without power correction is shown as a dashed line. For comparison, we also show the charged hadron data collected at the Z resonance by the DELPHI collaboration [22].

is shifted to the right by an amount O_{NP} and second, the distribution is rescaled by a factor $(O + O_{\text{NP}})/O$. In the region of the turnover where O is of the same order as O_{NP} there is an enhancement of $\mathcal{O}(100\%)$. The hadronisation correction is smaller at larger values of O .

For comparison, we also show the charged hadron data collected by the DELPHI Collaboration [22] at the Z resonance. We see remarkable agreement (strikingly so in view of the simplified hadronisation correction applied here). The only discrepancy is at very small values of $O < O_{\text{NP}}$ where individual hadrons in the light/narrow hemisphere will significantly affect the value of O . We do not expect to successfully describe such events.

6. Conclusions

In this paper, we presented predictions for the light hemisphere mass and the narrow jet broadening distributions where the infrared logarithms have been resummed to next-to-next-to-leading logarithmic order using the coherent branching formalism. The resummed expressions do not exponentiate, but involve the difference of exponential factors so that the perturbative series starts at $\mathcal{O}(\alpha_s^2)$. The expressions presented in sections 3 and 4 resum the leading $\mathcal{O}(\alpha_s^n L^{2n-1})$ and sub-leading infrared logarithms down to $\mathcal{O}(\alpha_s^n L^{2n-3})$ to all orders in the coupling constant. Taken in conjunction with the fixed order $\mathcal{O}(\alpha_s^3)$ non-logarithmic terms, the first neglected

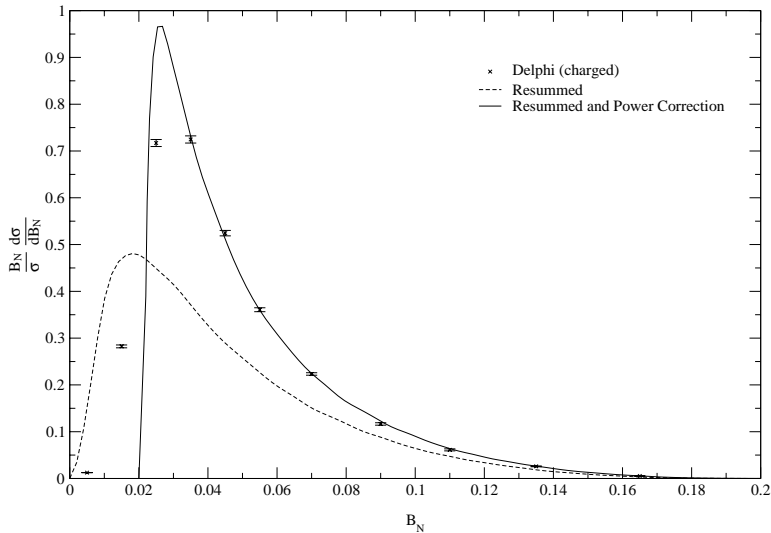


Figure 6: The resummed prediction for the narrow jet broadening distribution $\frac{B_N}{\sigma} \frac{d\sigma}{dB_N}$ at $\mu = Q = M_Z$ modified by a non-perturbative power correction $B_N \rightarrow B_N + 0.3 \text{ GeV} \ln(1/B_N)/Q$. The resummed prediction without power correction is shown as a dashed line. For comparison, we also show the charged hadron data collected at the Z resonance by the DELPHI collaboration [22].

term is of $\mathcal{O}(\alpha_s^4 L^5)$. The numerical results shown in Figs. 3 and 4 indicate that the resummation effects are sizeable at small values of the event shape parameter and that the resummation procedure significantly improves the perturbative prediction.

However, because important non-perturbative hadronisation corrections are present, resummation alone is insufficient to describe the experimental data. Applying simple power corrections similar to those obtained for the heavy hemisphere mass and the wide jet broadening gives good qualitative agreement with the available data from LEP. We anticipate that the improved theoretical description of the four-jet event shape distributions presented here can be combined with a more sophisticated hadronisation correction based on the dispersive approach of Ref. [20]. This will yield a theoretical description of the ρ_L and B_N distributions that is on a similar footing with the well studied three-jet event shape variables, $1 - T$, ρ_H , B_W etc. The data from LEP can then be used to further test the structure of QCD in four-jet like events and extract values of the strong coupling constant.

Acknowledgements

We thank Mike Seymour and David Summers for useful and stimulating discussions. SJB thanks the UK Particle Physics and Astronomy Research Council for research studentships. This work was supported in part by the EU Fourth Framework Programme ‘Training and Mobility of Researchers’, Network ‘Quantum Chromodynam-

ics and the Deep Structure of Elementary Particles', contract FMRX-CT98-0194 (DG-12-MIHT).

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